

P1

a) Nos piden normalizar la función de onda, esto significa imponer lo siguiente: $\langle \Psi | \Psi \rangle = 1 = \int_{-\infty}^{\infty} \langle \Psi | x \rangle \langle x | \Psi \rangle dx = \int_{-\infty}^{\infty} \Psi^*(x, 0) \Psi(x, 0) dx$
 $= \int_{-\infty}^{\infty} \frac{A^2}{(x^2 + a^2)^2} dx = 1$

Nos queda por resolver la integral, usamos: $x = a \tan u \Rightarrow dx = a \sec^2 u du$.

$$I = \int_{-\pi/2}^{\pi/2} \frac{a \sec^2 u}{(a^2 \tan^2 u + a^2)^2} du = \frac{a}{a^4} \int_{-\pi/2}^{\pi/2} \frac{\sec^2 u}{\sec^4 u} du = \frac{1}{a^3} \int_{-\pi/2}^{\pi/2} \cos^2 u du.$$

pero $\cos^2 u - \sin^2 u = \cos(2u) \Leftrightarrow \cos^2 u = \cos 2u + (1 - \cos^2 u) \Leftrightarrow 2\cos^2 u = 1 + \cos 2u$

$\Rightarrow \cos^2 u = \frac{1}{2}(1 + \cos 2u)$ reemplazamos en la integral: $\frac{A^2}{a^3} \left[\int_{-\pi/2}^{\pi/2} \frac{1}{2} du + \int_{-\pi/2}^{\pi/2} \cos(2u) du \right] \stackrel{!}{=} 1$

$$\Rightarrow A = \sqrt{\frac{2a^3}{\pi}}$$

(b) Veremos como se calcula cada uno de los términos solicitados:

i.- $\langle x \rangle = \langle \Psi | x | \Psi \rangle = \int_{-\infty}^{\infty} \Psi^*(x, 0) \times x \Psi(x, 0) dx = \int_{-\infty}^{\infty} \frac{A^2}{(x^2 + a^2)^2} x dx$ esto corresponde a una integral impar en dominio simétrico

$$\langle x \rangle = 0$$

ii) $\langle x^2 \rangle = \int_{-\infty}^{\infty} \frac{A^2 x^2}{(x^2 + a^2)^2} dx = A^2 \int_{-\infty}^{\infty} \frac{x^2 + a^2}{(x^2 + a^2)^2} - \frac{a^2}{(x^2 + a^2)^2} dx = A^2 \left[\arctan\left(\frac{x}{a}\right) \cdot \frac{1}{a} \right]_{-\infty}^{\infty} -$

$$\int_{-\infty}^{\infty} \frac{1}{(x^2 + a^2)^2} dx \Big] = \dots = \frac{A^2 \pi}{a^2}$$
 para calcular la segunda integral se usa $\int_{-\infty}^{\infty} \frac{1}{(ax^2 + b)^n} dx = \frac{2n-3}{2b(n-1)} \int_{-\infty}^{\infty} \frac{1}{(ax^2 + b)^{n-1}} dx + \frac{x}{2b(n-1)(ax^2 + b)^{n-1}}$

iii) $\Delta x \equiv \sqrt{\langle x^2 \rangle - \langle x \rangle^2} = \sqrt{\langle x^2 \rangle} = \sqrt{\frac{A^2 \pi}{a^2}}$

P2

$$\Psi(x, t) = A e^{-a\left(\frac{mx^2}{\hbar} + it\right)}$$

a) $\langle \Psi | \Psi \rangle \stackrel{!}{=} 1 = \int_{-\infty}^{\infty} \Psi^*(x, t) \Psi(x, t) dx = \int_{-\infty}^{\infty} |A|^2 e^{-2a\frac{mx^2}{\hbar}} dx = A^2 \int_{-\infty}^{\infty} e^{-x^2 / \sqrt{\hbar/2ma}} dx$

integral que tiene la forma gaussiana que sabemos $\int_{-\infty}^{\infty} e^{-\frac{x^2}{\sigma^2}} dx = \sqrt{\pi \sigma^2}$ con $\sigma = \sqrt{\frac{\hbar}{2ma}}$

$$\Rightarrow 1 = A^2 \sqrt{\frac{\pi \hbar}{2am}} \Rightarrow A = \left(\frac{\pi \hbar}{2am}\right)^{\frac{1}{4}} //$$

b) Queremos $\phi(k, 0)$, para eso debemos calcular la transformada de Fourier de $\psi(x, 0)$:

$$\begin{aligned} \phi(k) &= \int_{-\infty}^{\infty} A e^{-\frac{amx^2}{\hbar}} e^{-ikx} dx = A \int_{-\infty}^{\infty} e^{-\left(\frac{x^2}{\sigma^2} + kx\right)} dx = A \int_{-\infty}^{\infty} e^{-\left(\frac{x}{\sigma} + \frac{\sigma}{2}k\right)^2 - \frac{\sigma^2}{4}k^2} dx \\ &= A e^{-\frac{\sigma^2}{4}k^2} \int_{-\infty}^{\infty} e^{-\left(\frac{x+\frac{\sigma}{2}k}{\sigma}\right)^2} dx = A e^{-\frac{\sigma^2}{4}k^2} \sqrt{\pi} \sigma // \end{aligned}$$

ojo! A es el mismo que encontramos en (a), no es necesario normalizar de nuevo

c) $\langle x \rangle = \int_{-\infty}^{\infty} x |\psi|^2 dx = \int_{-\infty}^{\infty} x A^2 e^{-\frac{2amx^2}{\hbar}} dx = 0$ integrando impar en dominio simétrico.

$$\langle x^2 \rangle = \int_{-\infty}^{\infty} x^2 |\psi|^2 dx = A^2 \int_{-\infty}^{\infty} x^2 e^{-\frac{2amx^2}{\hbar}} dx$$

Γ truco: $\int_{-\infty}^{\infty} e^{-ax^2} dx = \sqrt{\frac{\pi}{a}} \Rightarrow \int_{-\infty}^{\infty} -x^2 e^{-ax^2} dx = \sqrt{\pi} \cdot \left(-\frac{1}{2}\right) a^{-3/2} \Rightarrow \int_{-\infty}^{\infty} x^2 e^{-ax^2} dx = \frac{1}{2} \sqrt{\frac{\pi}{a^3}}$

$$\langle x^2 \rangle = A^2 \cdot \frac{1}{2} \sqrt{\pi \left(\frac{\hbar}{2am}\right)^3} = A^2 \cdot \frac{1}{2} \frac{1}{(2am/\hbar)} \sqrt{\frac{\pi \hbar}{2am}} = \frac{\hbar}{4am} //$$

$$\langle p \rangle = m \frac{d\langle x \rangle}{dt} = 0 ; \text{ si no tuviese la función en el tiempo esto no se puede hacer, habría que calcular } \langle \psi | p | \psi \rangle = \langle \psi | -i\hbar \frac{d}{dx} | \psi \rangle$$

que calcular $\langle \psi | p | \psi \rangle = \langle \psi | -i\hbar \frac{d}{dx} | \psi \rangle$

$$\begin{aligned} \langle p^2 \rangle &= \int \psi^* \left(\frac{\hbar}{i} \frac{d}{dx}\right)^2 \psi dx = -\hbar^2 \int \psi^* \frac{\partial^2 \psi}{\partial x^2} dx = -\hbar^2 \int_{-\infty}^{\infty} \psi^* \left[-\frac{2am}{\hbar} \left(1 - \frac{2amx^2}{\hbar}\right) \psi\right] dx \\ &= 2am\hbar \left[\int |\psi|^2 dx - \frac{2am}{\hbar} \int x^2 |\psi|^2 dx \right] = 2am\hbar \left(1 - \frac{2am}{\hbar} \langle x^2 \rangle\right) = 2am\hbar \left(1 - \frac{2am}{\hbar} \cdot \frac{\hbar}{4am}\right) = \dots = am\hbar \end{aligned}$$

d) $\sigma_x^2 = \Delta x^2 = \langle x^2 \rangle - \langle x \rangle^2 = \frac{\hbar}{4am} \Rightarrow \sigma_x = \sqrt{\frac{\hbar}{4am}}$
 $\sigma_p^2 = \langle p^2 \rangle - \langle p \rangle^2 = am\hbar \Rightarrow \sigma_p = \sqrt{am\hbar}$

} $\sigma_x \sigma_p = \sqrt{\frac{\hbar}{4ma}} \cdot \sqrt{am\hbar} = \frac{\hbar}{2} //$ cumple! justito...

P3

$$\hbar / \sqrt{3mk_B T} > d \Rightarrow T < \frac{\hbar^2}{3mk_B d^2}$$

a) electrones ($m_e = 9.1 \cdot 10^{-31} \text{ kg}$): $T < \frac{(6.6 \cdot 10^{-34})^2}{3(9.1 \cdot 10^{-31})(1.4 \cdot 10^{-23})(3 \cdot 10^{-9})^2} = 1.3 \cdot 10^5 \text{ K}$

núcleos de Silicio ($m = 28 m_p = 28(1.7 \cdot 10^{-27}) = 4.76 \cdot 10^{-26} \text{ kg}$)

$$T < \frac{(6.6 \cdot 10^{-34})^2}{3(4.76 \cdot 10^{-26})(1.4 \cdot 10^{-23})(3 \cdot 10^{-9})^2} \approx 2.4 \text{ K}$$

b) $PV = Nk_B T$; volumen ocupado por una molécula ($N=1, V=d^3 \Rightarrow d = (k_B T / P)^{1/3}$)

$$T < \frac{\hbar^2}{2mk_B} \left(\frac{P}{k_B T}\right)^{2/3} \Rightarrow T^{5/2} = \frac{\hbar^2}{2m} \frac{P^{2/3}}{k_B^{5/3}} \Rightarrow T < \frac{1}{k_B} \left(\frac{\hbar^2}{2m}\right)^{3/5} P^{2/5}$$

para el helio ($m = 4m_p = 6.8 \cdot 10^{-27} \text{ kg}$) a $1 \text{ atm} = 1 \cdot 10^5 \frac{\text{N}}{\text{m}^2}$

$$T < \frac{1}{(1.4 \cdot 10^{-23})} \cdot \left(\frac{(6.6 \cdot 10^{-34})^2}{3(6.8 \cdot 10^{-27})} \right)^{2/5} \cdot (10^5)^{2/5} = 2.8 \text{ K}$$

para el hidrogeno ($m = 2m_p = 3.4 \cdot 10^{-27} \text{ kg}$) con $d = 10^{-2} \text{ m}$

$$T < \frac{(6.6 \cdot 10^{-34})^2}{3(3.4 \cdot 10^{-27})(1.4 \cdot 10^{-23})(10^{-2})^2} = 3.1 \cdot 10^{14} \text{ K}$$

Si el hidrogeno espacial está a 3 K constante entonces es definitivamente en el regimen clásico.