

ÓRBITAS PLANETARIAS

$$g = g_0 \frac{1+e}{1+e \cos \theta} \quad (\theta = 0)$$

$$g_0 = \frac{h^2}{GM(1+e)}$$

$$h = \frac{b_0}{m}$$

SISTEMA SOLAR

EXCENTRICIDADES (e)

TIERRA = 0.017

MARTE : 0.093

VENUS ; 0.007

JÚPITER : 0.048

PLUTÓN : 0.249

COMETA HALLEY : 0.967

ÓRBITAS SATELITALES TERRESTRES

$$\frac{GM}{R_T^2} = g$$

$$g_0 = \frac{h^2}{g R_T (1+e)}$$

ÓRBITA GEOESTACIONARIA ECUATORIAL

EL SATELITE SE MUEVE CON LA MISMA VELOCIDAD ANGULAR QUE LA TIERRA

$$\omega_0 = \frac{2\pi}{T}$$

$$T = 24 \times 3600$$

$$T = 86400$$

$$m(\ddot{r} - r\dot{\theta}^2) = -\frac{GM_T m}{r^2} = -\frac{gR_T^2 m}{r^2}$$

$\downarrow$
 ω_0^2

$$r^3 = \frac{gR_T^2}{\omega_0^2}$$

$$R_T = 6360 \text{ km}$$

$$= 6.36 \times 10^6 \text{ m}$$

$$r^3 = \frac{9.81 \cdot 6.36^2 \cdot 10^{12} \cdot 8.64^2 \cdot 10^8}{4\pi^2}$$

$$r = 42.177 \text{ km}$$

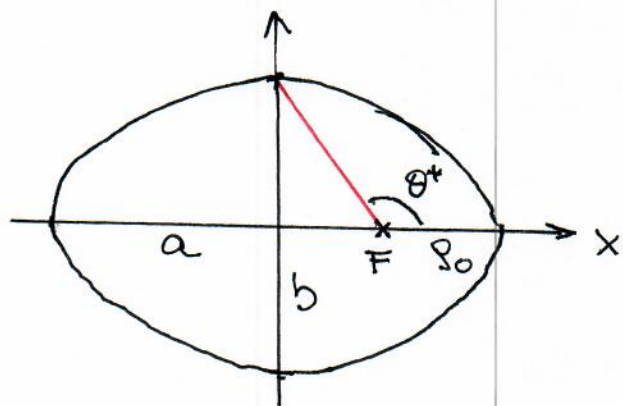
ALTURA DE LA ÓRBITA GEOSTACIONARIA
SOBRE LA SUPERFICIE DE LA TIERRA

$$h \quad \left| \quad z^* = 35.817 \text{ km} \quad \right|$$

ÓRBITA ELÍPTICA

TERCERA LEY DE KEPLER

$$(\text{PERIODO ORBITAL})^2 \propto (\text{SEMI-EJE MAYOR})^3$$



$$a = \frac{1}{2}(r_0 + r_1)$$

$$r_0 = r(\theta=0)$$

$$r_1 = r(\theta=\pi)$$

$$r = r_0 \frac{1+e}{1+e\cos\theta} \rightarrow r_1 = r_0 \frac{(1+e)}{(1-e)}$$

$$a = \frac{1}{2} \left[r_0 + \frac{r_0(1+e)}{(1-e)} \right] = \frac{r_0}{1-e} \quad \left| a = \frac{r_0}{1-e} \right|$$

$$y = r \sin\theta \quad y_{\text{MAX}} = b$$

$$y = r_0(1+e) \underbrace{(1+e\cos\theta)^{-1}}_{f(\theta)} \sin\theta$$

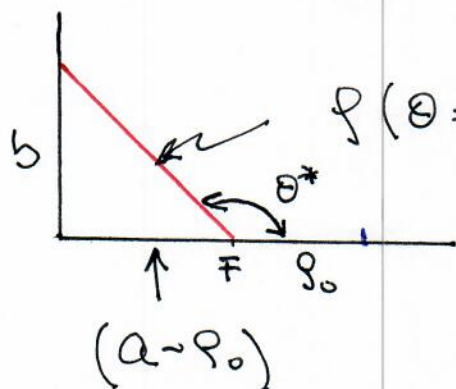
HAY QUE ENCONTRAR EL MAX. DE $f(\theta)$

$$\frac{df}{d\theta} = 0 \rightarrow -(1+e\cos\theta)^{-2} \overset{e}{\downarrow} (-\sin\theta) \sin\theta + (1+e\cos\theta)^{-1} \cos\theta$$

$$\frac{df}{d\theta} = \frac{e\sin^2\theta + (1+e\cos\theta)\cos\theta}{(1+e\cos\theta)^2} = 0$$

$$\frac{df}{d\theta} = 0 \rightarrow e + \cos\theta = 0$$

$$\cos\theta^* = -e$$



$$f(\theta = \theta^*) = \frac{p_0 (1 + e)}{1 + e \cos\theta^*}$$

$$= \frac{p_0 (1 + e)}{1 - e^2} = \frac{p_0}{1 - e} = a$$

$$a - p_0 = \frac{p_0}{1 - e} - p_0 = \frac{p_0 e}{1 - e} = ae$$

$$b^2 + (a - p_0)^2 = f^2(\theta^*)$$

$$b = \sqrt{a^2 - (ae)^2} \rightarrow b = a\sqrt{1 - e^2}$$

2^o LEY DE KEPLER

$$\frac{ds}{dt} = \frac{1}{2m} l_0$$

$$S = \frac{1}{2m} l_0 T$$

T: PERIODO ORBITAL

l_0 : MAGNITUD DEL
MOMENTUM
ANGULAR
CONSTANTE

EN EL CASO DE UNA ÓRBITA ELÍPTICA (5)

$$| S = \pi a b |$$

$$T = \frac{2m\pi ab}{l_0} = \frac{2\pi ab}{h} \quad h = \frac{l_0}{m}$$

$$T^2 = \frac{(4\pi)^2}{h^2} a^2 b^2 = \frac{(4\pi)^2}{h^2} a^2 a^2 (1-e^2)$$

$$a(1-e^2) = \frac{p_0}{1-e} (1-e^2) = p_0(1+e)$$

$$\text{PERO } p_0(1+e) = \frac{h^2}{GM}$$

$$T^2 = \frac{(4\pi)^2}{h^2} a^3 \frac{h^2}{GM}$$

$$| T^2 = \frac{(4\pi)^2}{GM} a^3 |$$

ENERGÍA MECÁNICA Y EXCENTRICIDAD

FUERZA GRAVITACIONAL

$$\vec{F} = - \frac{GMm}{r^2} \hat{r}$$

POTENCIAL GRAVITATORIO

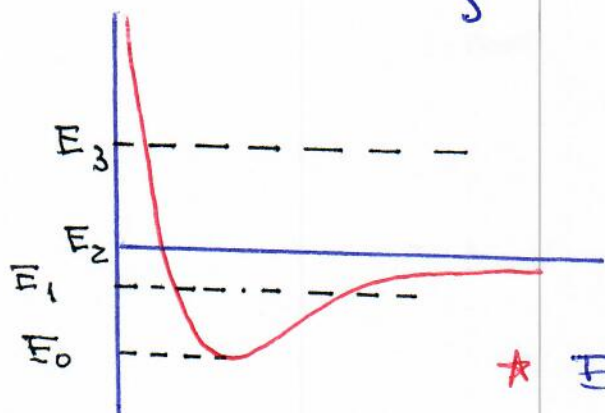
$$V = - \frac{GMm}{r}$$

$$V^* = - \frac{GMm}{r} + \frac{m h^2}{2 r^2}$$

$$V = 0 \quad r = \infty$$

EC. CONSERVACIÓN ENERGÍA

$$\left| \frac{1}{2} m \dot{r}^2 + V^* = E_0 \right|$$



★ ENERGÍA E_0 (< 0) ÓRBITA CIRCULAR

★ ENERGÍA E_1 (< 0) ÓRBITA ELÍPTICA

★ ENERGÍA E_2 ($= 0$) ÓRBITA PARABÓLICA

$$\lim_{r \rightarrow \infty} K_r = 0$$

★ ENERGÍA E_3 (> 0) ÓRBITA HIPERBÓLICA

$$\lim_{r \rightarrow \infty} K_r = E_3$$

$$\left| K_r = \frac{1}{2} m \dot{r}^2 \right|$$

$$\frac{1}{2} m v^2 + V(r) = E_0$$

$$v^2 = \dot{r}^2 + (r\dot{\theta})^2$$

$$\xi = \frac{1}{r}$$

$$\frac{dr}{dt} = \frac{dr}{d\theta} \dot{\theta} = \frac{d}{d\theta} \left(\frac{1}{\xi} \right) \dot{\theta} \quad \left| \quad \overline{r^2 \dot{\theta} = \frac{L_0}{m} = h} \right.$$

$$\frac{dr}{dt} = - \frac{1}{\cancel{\xi^2}} \frac{d\xi}{d\theta} \cdot \cancel{h \xi^2} = -h \left(\frac{d\xi}{d\theta} \right) \quad \left| \quad \overline{\dot{\theta} = h \xi^2} \right.$$

$$v^2 = h^2 \left(\frac{d\xi}{d\theta} \right)^2 + \frac{1}{\xi^2} h^2 \xi^4$$

$$v^2 = h^2 \left[\left(\frac{d\xi}{d\theta} \right)^2 + \xi^2 \right]$$

$$E = \frac{m}{2} h^2 \left[\left(\frac{d\xi}{d\theta} \right)^2 + \xi^2 \right] - GMm\xi$$

$$\frac{2E}{m} = h^2 \left[\left(\frac{d\xi}{d\theta} \right)^2 + \xi^2 \right] - 2GM\xi$$

PERO

$$\xi = \frac{1 + e \cos \theta}{r_0 (1 + e)}$$

⑧

$$\frac{df}{d\theta} = \frac{-e \sin \theta}{p_0(1+e)}$$

$$\frac{2E}{m} = h^2 \left[\left(\frac{-e \sin \theta}{p_0(1+e)} \right)^2 + \frac{(1+e \cos \theta)^2}{p_0^2(1+e)^2} \right] - 2GM \frac{1+e \cos \theta}{p_0(1+e)}$$

$$= h^2 \frac{(e^2 + 1 + 2e \cos \theta)}{[p_0(1+e)]^2} - 2GM \frac{(1+e \cos \theta)}{p_0(1+e)}$$

PERO $\left| p_0(1+e) = \frac{h^2}{GM} \right|$

$$\frac{2E}{m} = \frac{h^2 (GM)^2}{h^4} (e^2 + 1 + 2e \cos \theta) - 2GM \cdot \frac{GM}{h^2} (1+e \cos \theta)$$

$$\frac{2E}{m} = \left(\frac{GM}{h} \right)^2 (e^2 - 1)$$

$$e = \left[1 + \frac{2E}{m} \left(\frac{h}{GM} \right)^2 \right]^{1/2}$$

$$E < 0 \rightarrow e < 1$$

CÍRCULO + ELIPSE

$$E = 0 \rightarrow e = 1$$

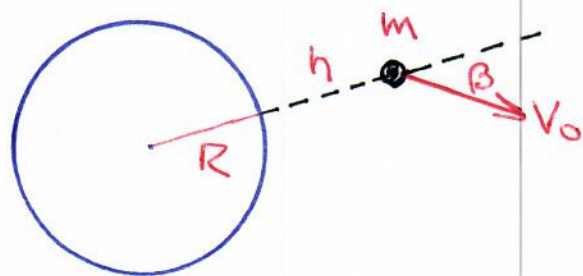
PARÁBOLA

$$E > 0 \rightarrow e > 1$$

HIPÉRBOLA

EJ. D.19

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¿ EC. DE TRAYECTORIA ?

$$E = \frac{1}{2} m V_0^2 - \frac{GMm}{(R+h)}$$

$$h = (R+h) V_0 \sin \beta$$

$$e = \left[1 + \frac{2E}{m} \left(\frac{h}{GM} \right)^2 \right]^{1/2}$$

$$p_0(1+e) = \frac{h^2}{GM} \rightarrow p_0 = \frac{h^2}{GM} \frac{1}{(1+e)}$$

ECUACIÓN DE TRAYECTORIA

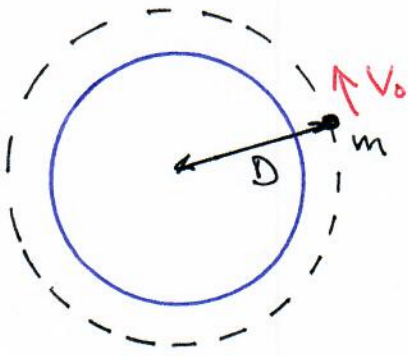
$$p = p_0 \frac{1+e}{1+e \cos \theta} = \frac{h^2}{GM} \frac{1}{1+e \cos \theta}$$

$$p(\theta) = \frac{h^2}{GM} \frac{1}{1+e \cos \theta}$$

EJ. D.21

SATÉLITE EN
ÓRBITA CIRCULAR DE RADIO D

(10)



$$\cancel{m} \frac{v_0^2}{D} = \frac{GM \cancel{m}}{D^2}$$

$$v_0 = \left(\frac{GM}{D} \right)^{1/2}$$

★ IMPULSO Δv TANGENCIAL PARA QUE EL SATÉLITE
ESCAPE DEL CAMPO DE ATRACCIÓN TERRESTRE

$$\frac{1}{2} m v^*{}^2 - \frac{GMm}{D} = 0 \rightarrow v^* = \left(\frac{2GM}{D} \right)^{1/2}$$

$$\Delta v = \left(\frac{2GM}{D} \right)^{1/2} - \left(\frac{GM}{D} \right)^{1/2} = \left(\frac{GM}{D} \right)^{1/2} (\sqrt{2} - 1)$$

★ IDEM CON UN Δv RADIAL

$$\frac{1}{2} \cancel{m} (v_0^2 + (\Delta v)^2) - \frac{GM \cancel{m}}{D} = 0$$

$$(\Delta v)^2 = 2 \frac{GM}{D} - v_0^2 = 2 \frac{GM}{D} - \frac{GM}{D}$$

$$\Delta v = \left(\frac{GM}{D} \right)^{1/2}$$